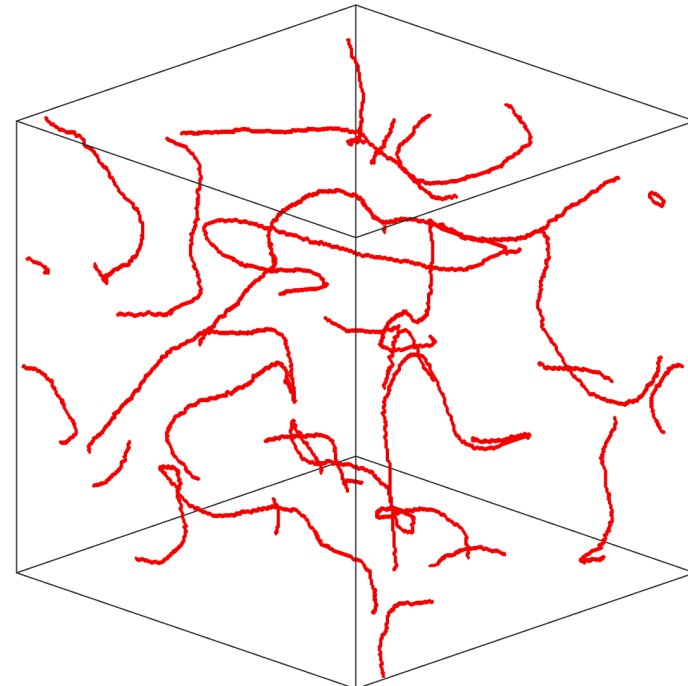
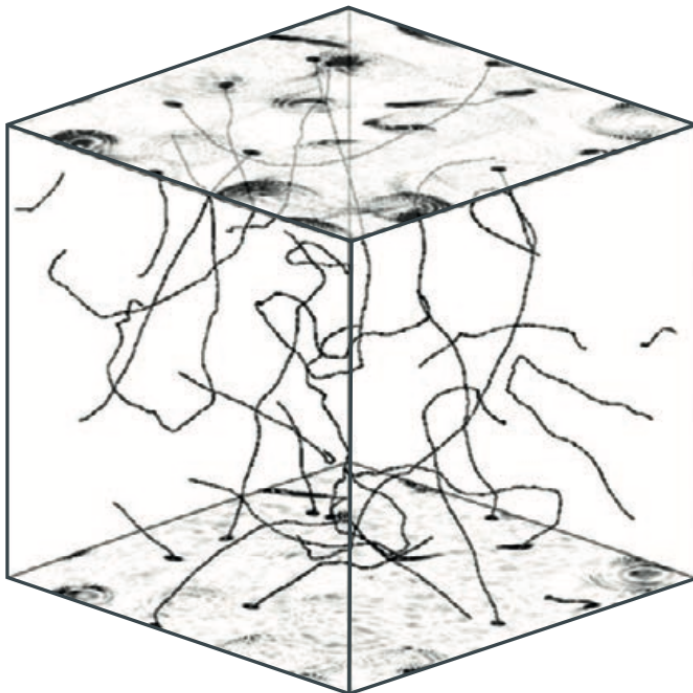
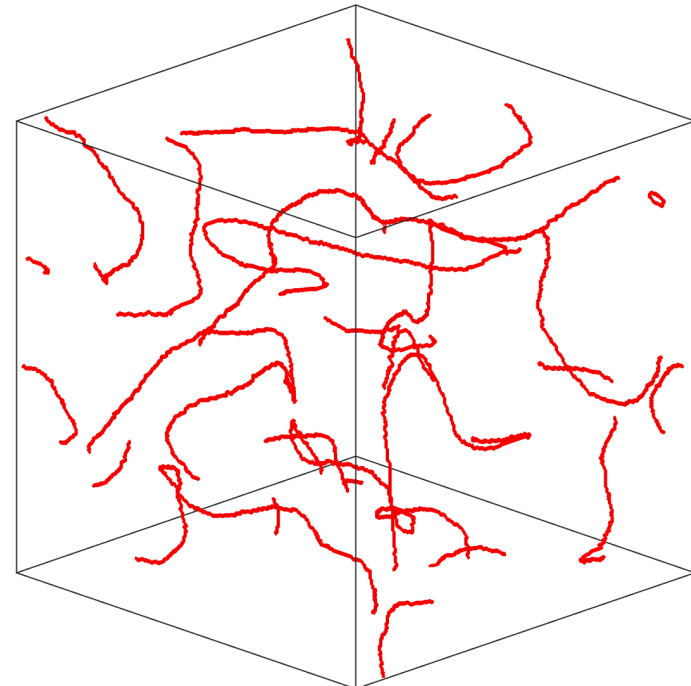


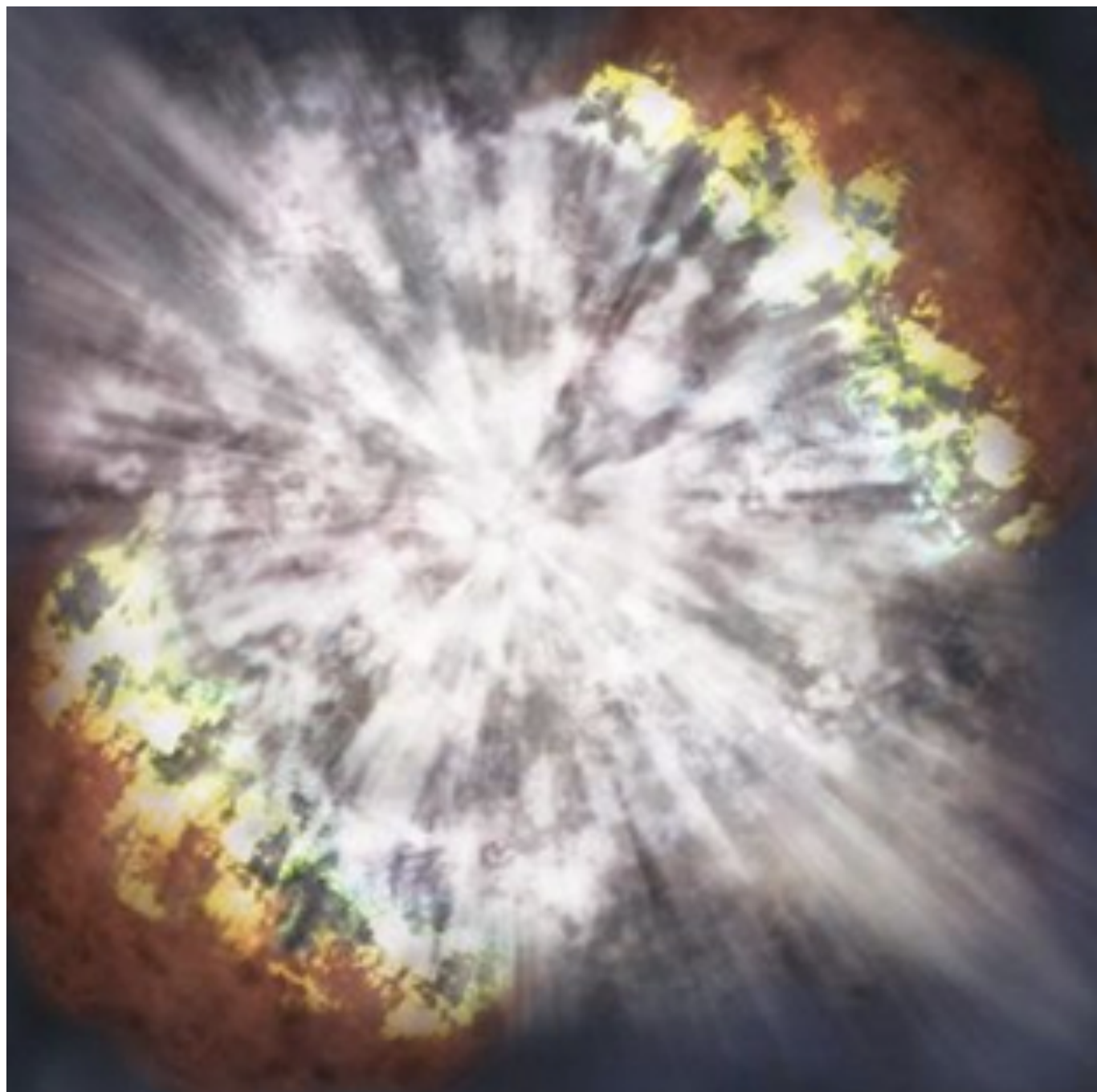
Computing strings, from the atomic to the cosmic

Nicola Spaldin, Materials Theory, ETH Zürich



Computing strings, from the atomic to the cosmic





HOMOGENEOUS VACUUM

small



After 10^{-37} seconds there is a spontaneous symmetry-lowering phase transition
the Grand Unification Transition (GUT)



larger

HOMOGENEOUS VACUUM

small

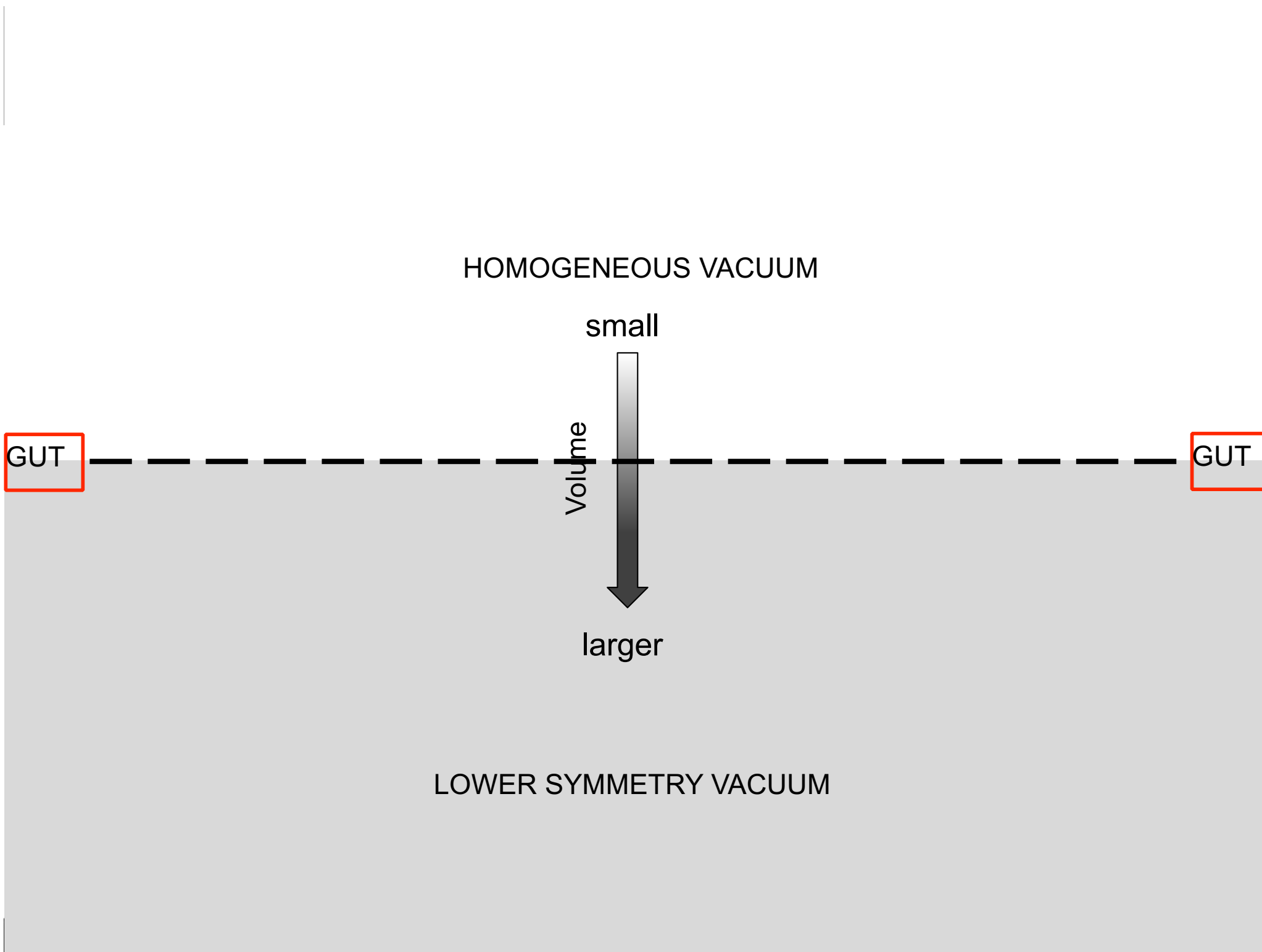
Volume

larger

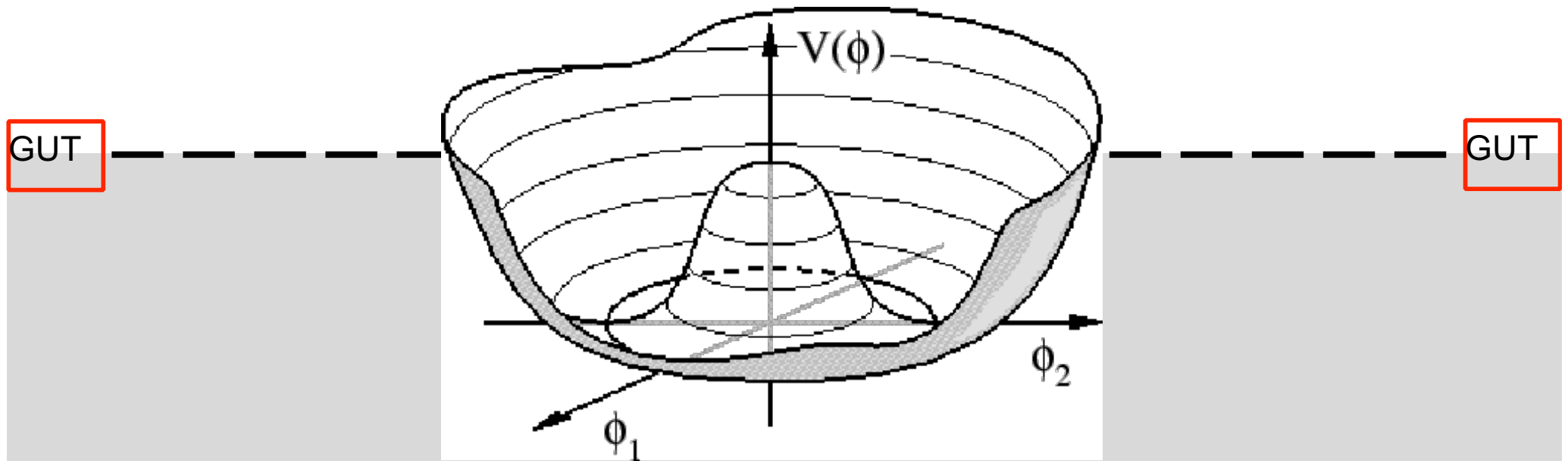
GUT

GUT

LOWER SYMMETRY VACUUM



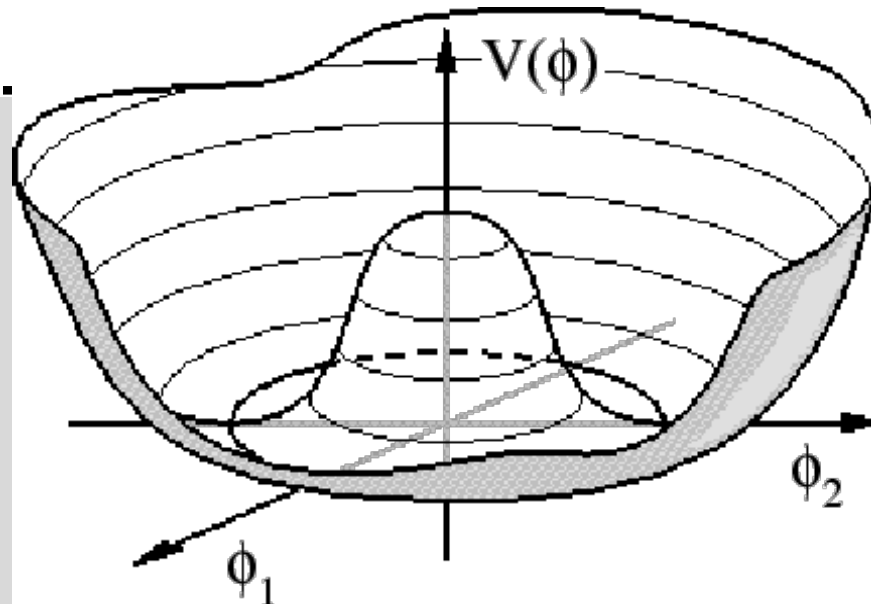
HOMOGENEOUS VACUUM



LOWER SYMMETRY VACUUM

Symmetry lowering at the Grand Unification Transition

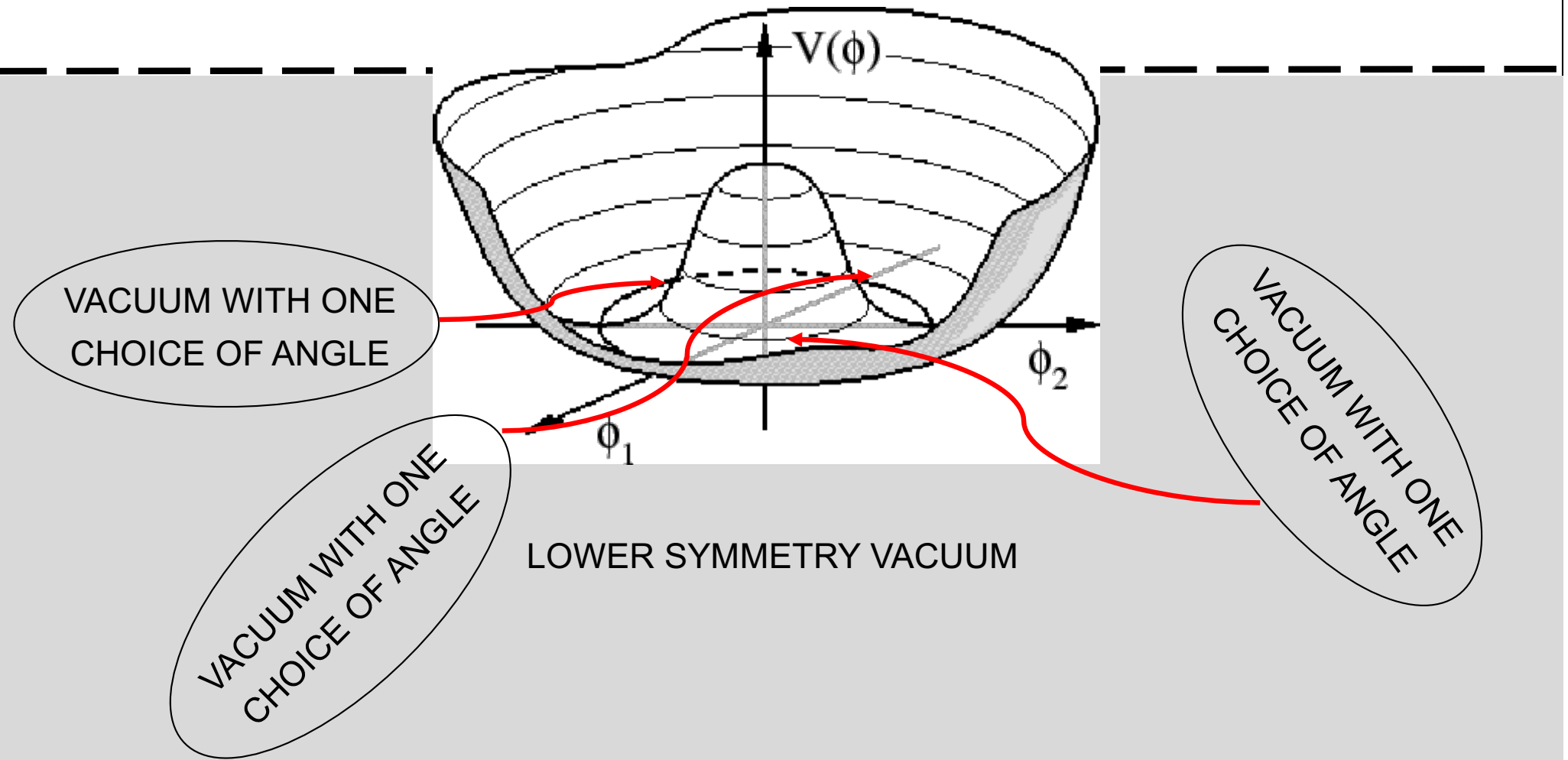
HOMOGENEOUS VACUUM



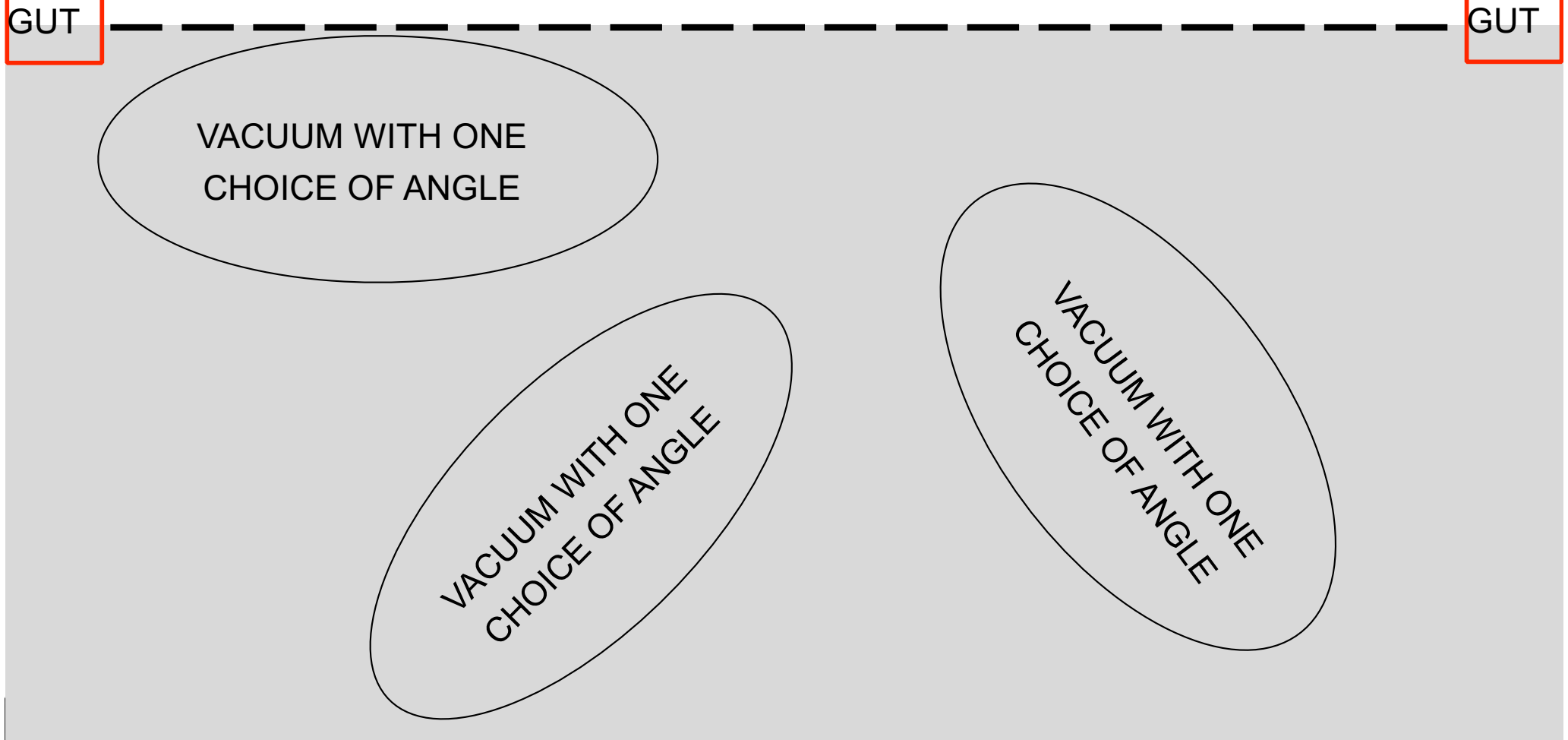
LOWER SYMMETRY VACUUM

Symmetry lowering at the Grand Unification Transition

HOMOGENEOUS VACUUM



As the universe expands through the transition, the low symmetry regions grow...



and grow...

GUT

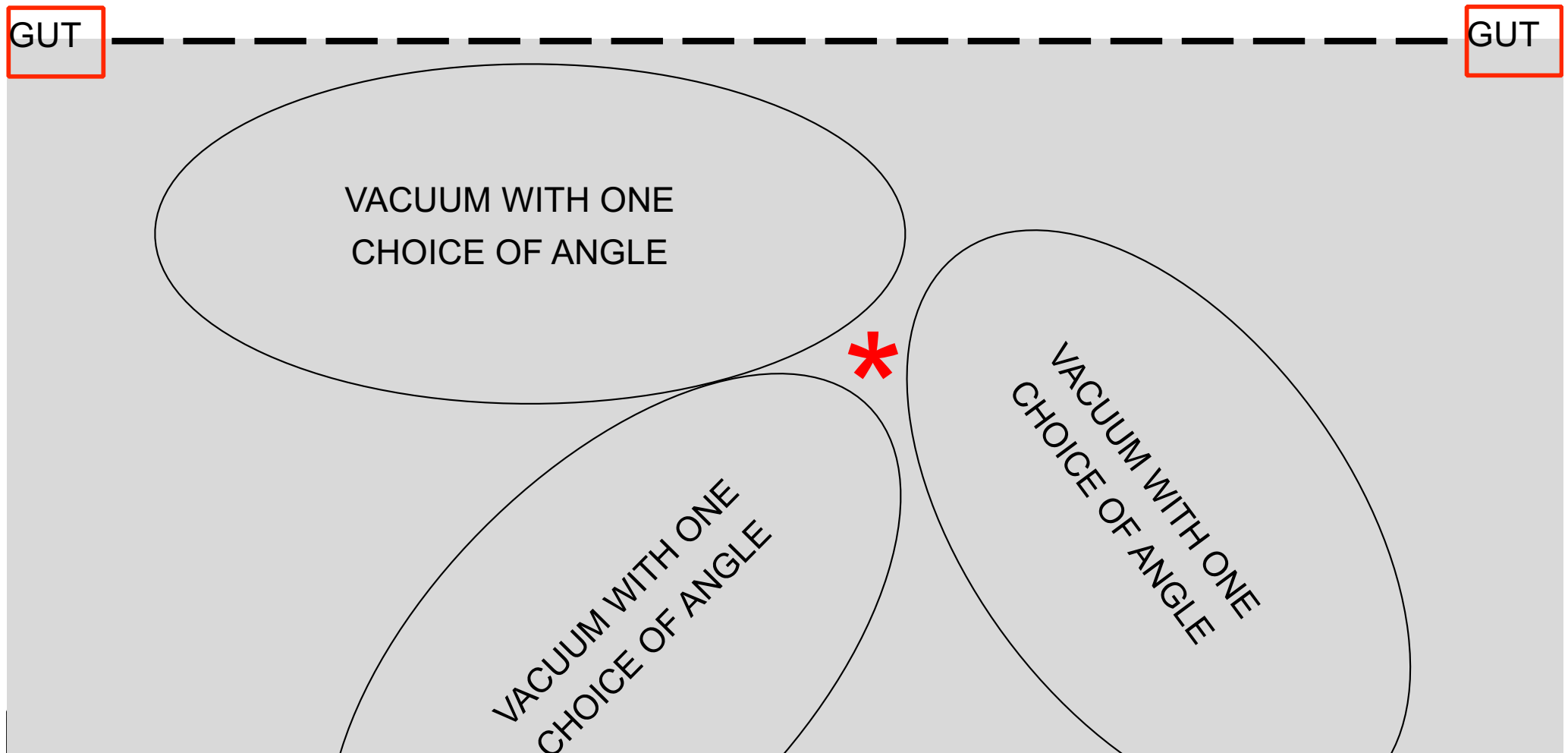
GUT

VACUUM WITH ONE
CHOICE OF ANGLE

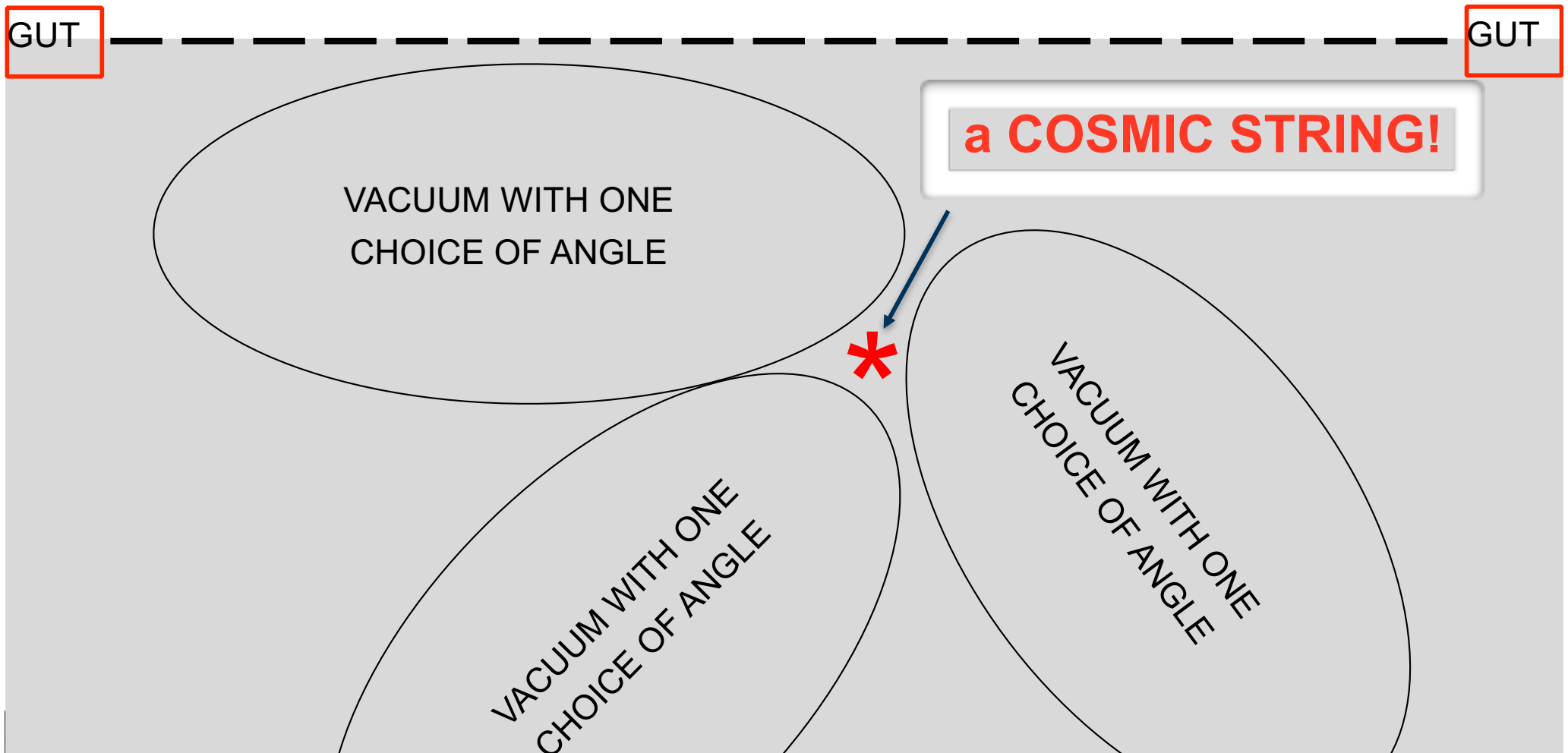
VACUUM WITH ONE
CHOICE OF ANGLE

VACUUM WITH ONE
CHOICE OF ANGLE

and eventually meet!



The angle mismatch in the vacuum is a *topologically protected one-dimensional defect* – a COSMIC STRING

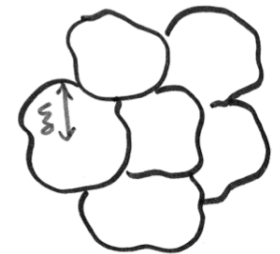


A detail: How many cosmic strings should we have?

It depends on the rate of expansion through the transition

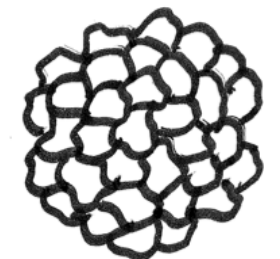
Expand slowly: Different regions can communicate their choice of angle

- **Large regions of the same choice**
- **Low density of cosmic strings**



Expand quickly: Not much time to communicate choice of angle

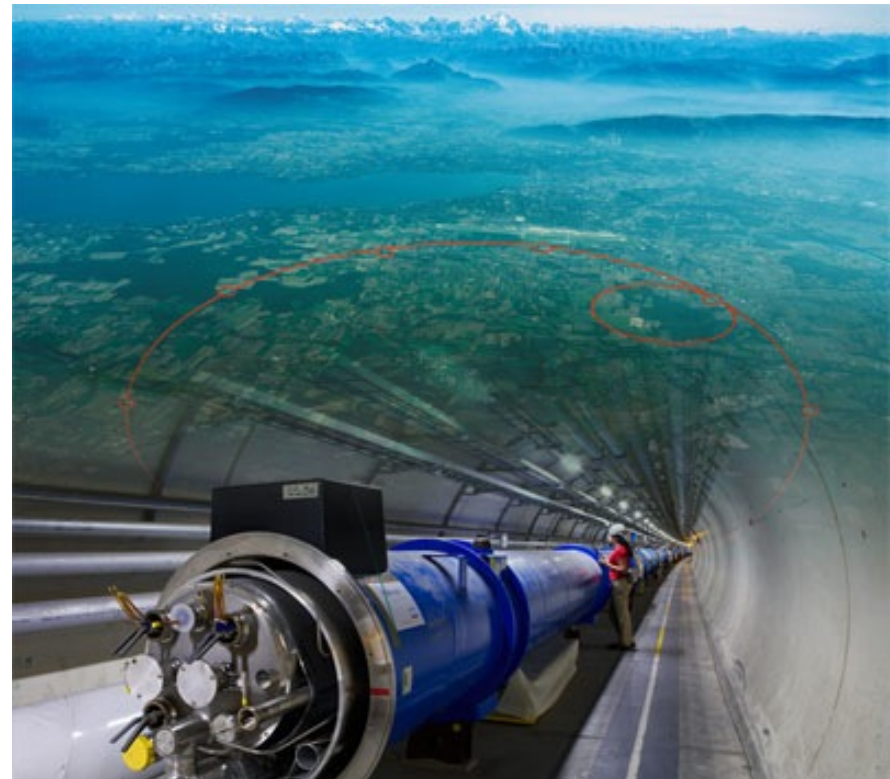
- **Many smaller regions with different choices of angle**
- **High density of cosmic strings**



Do cosmic strings exist? How can we study them?

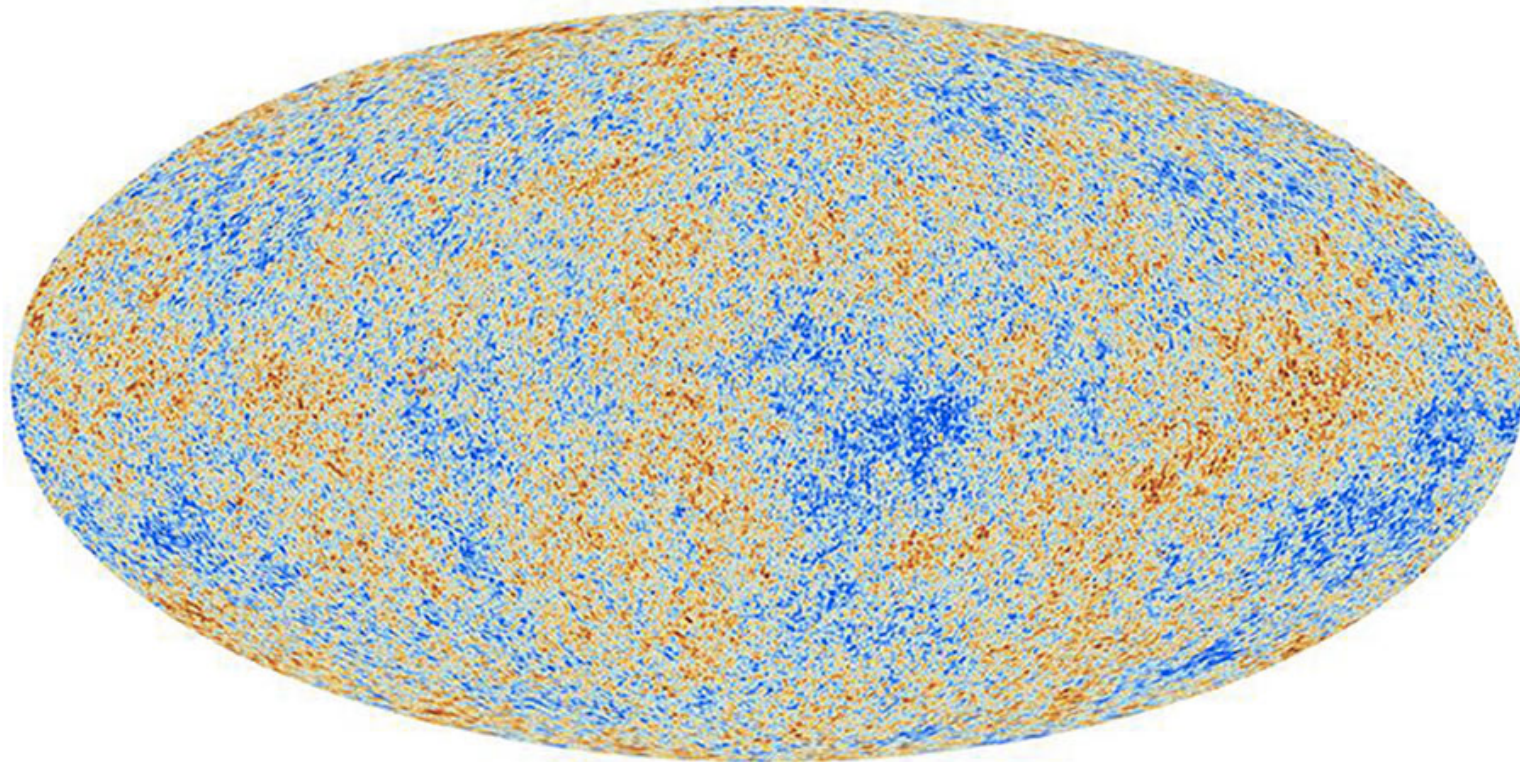
For direct study we need a probe with a similar energy, $\sim 10^{15}$ GeV

Our highest energy probes, the largest hadronic colliders reach $\sim 10,000$ GeV



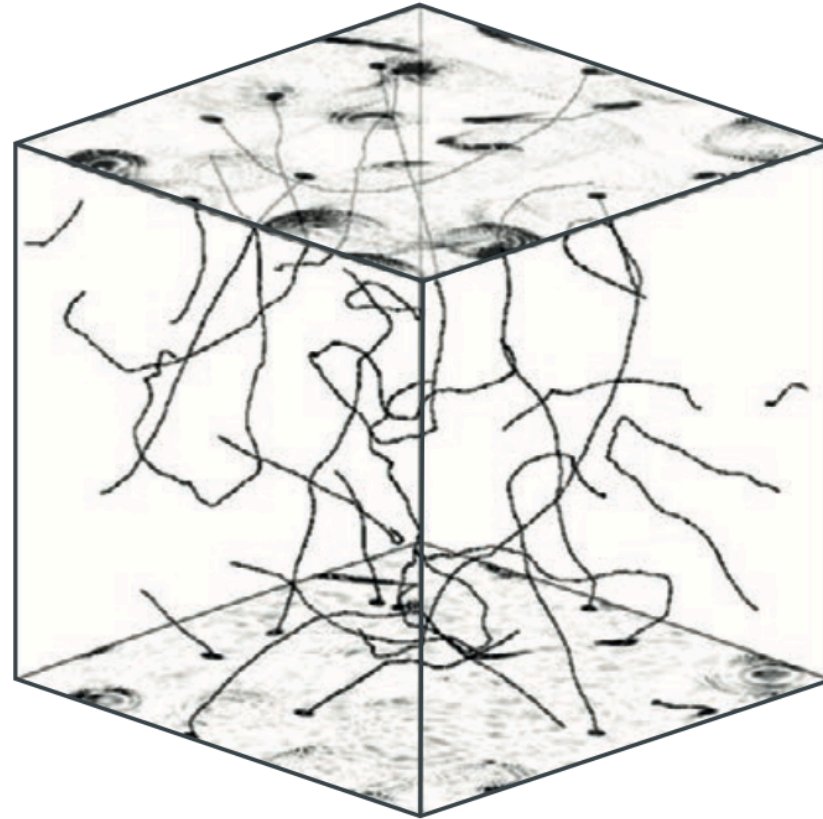
How is Cosmic String Formation at the Grand Unification Transition studied?

Analyzing the Cosmic Microwave Background



How is Cosmic String Formation at the Grand Unification Transition studied?

Computer Simulation



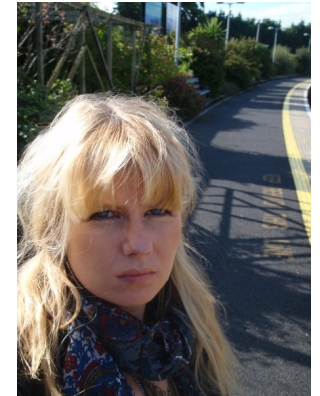
PHYSICAL REVIEW D **76**, 043005 (2007)

CMB polarization power spectra contributions from a network of cosmic strings

Neil Bevis,^{1,*} Mark Hindmarsh,^{1,†} Martin Kunz,^{2,‡} and Jon Urrestilla^{1,§}

Instead we will study the GUT in our laboratory!

First we will identify a material with a symmetry-lowering phase transition described by the same mathematics as that proposed for the GUT



Sinead Griffin

spontaneous symmetry breaking described by a Mexican hat potential

Then we will do experiments on the material to answer questions about the GUT:

Do cosmic strings exist?
Did they form as we think?
How did they evolve?
What are their properties?

Outline

Identify a material with a symmetry-lowering phase transition described by a Mexican-hat potential

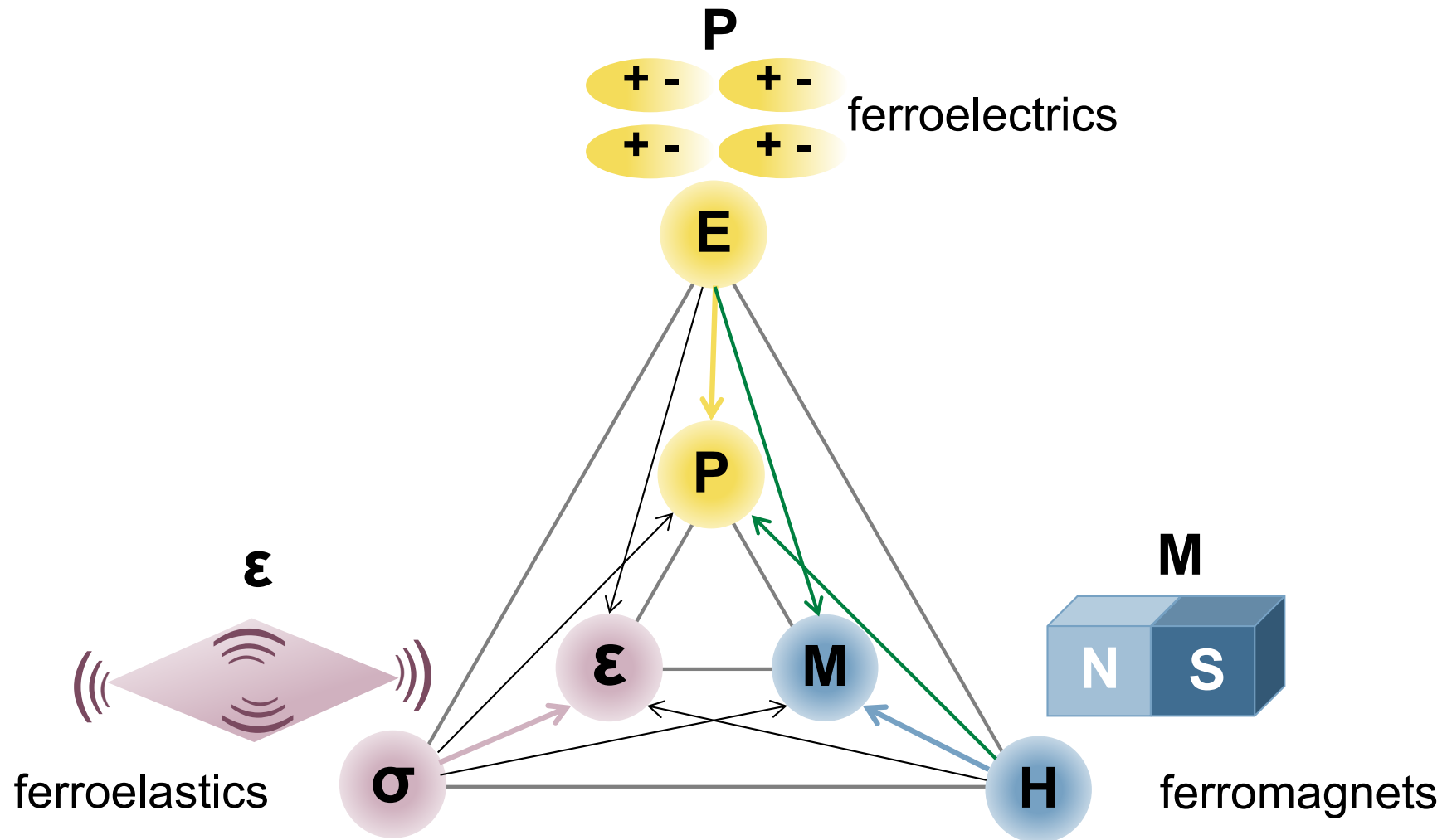
Use electronic structure calculations to calculate how many string-like defects should be formed as a function of cooling rate, based on the cosmic-string-formation model

Measure how many string-like defects are formed as a function of cooling rate

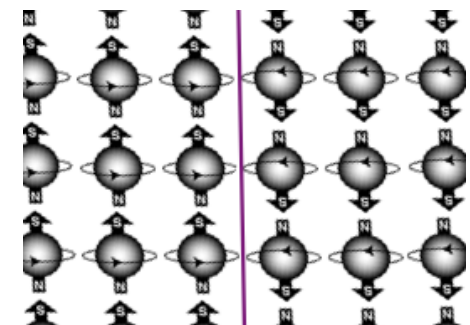
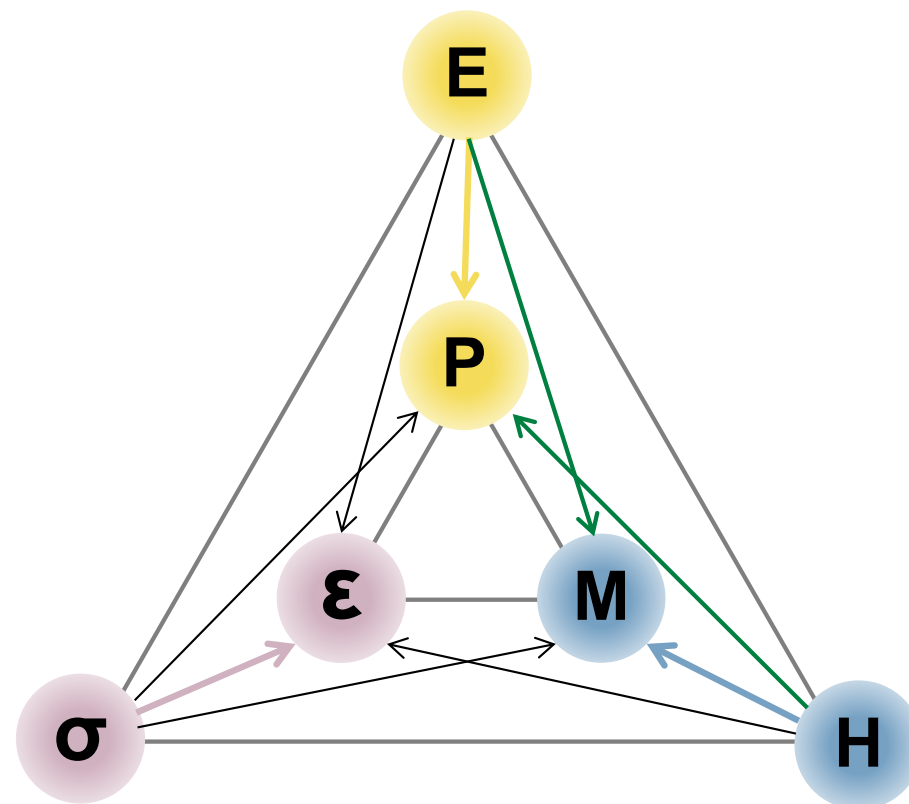
Does a system that is described by the same physics and symmetry as the GUT exhibit the predicted behavior?!

Where to find a suitable material?

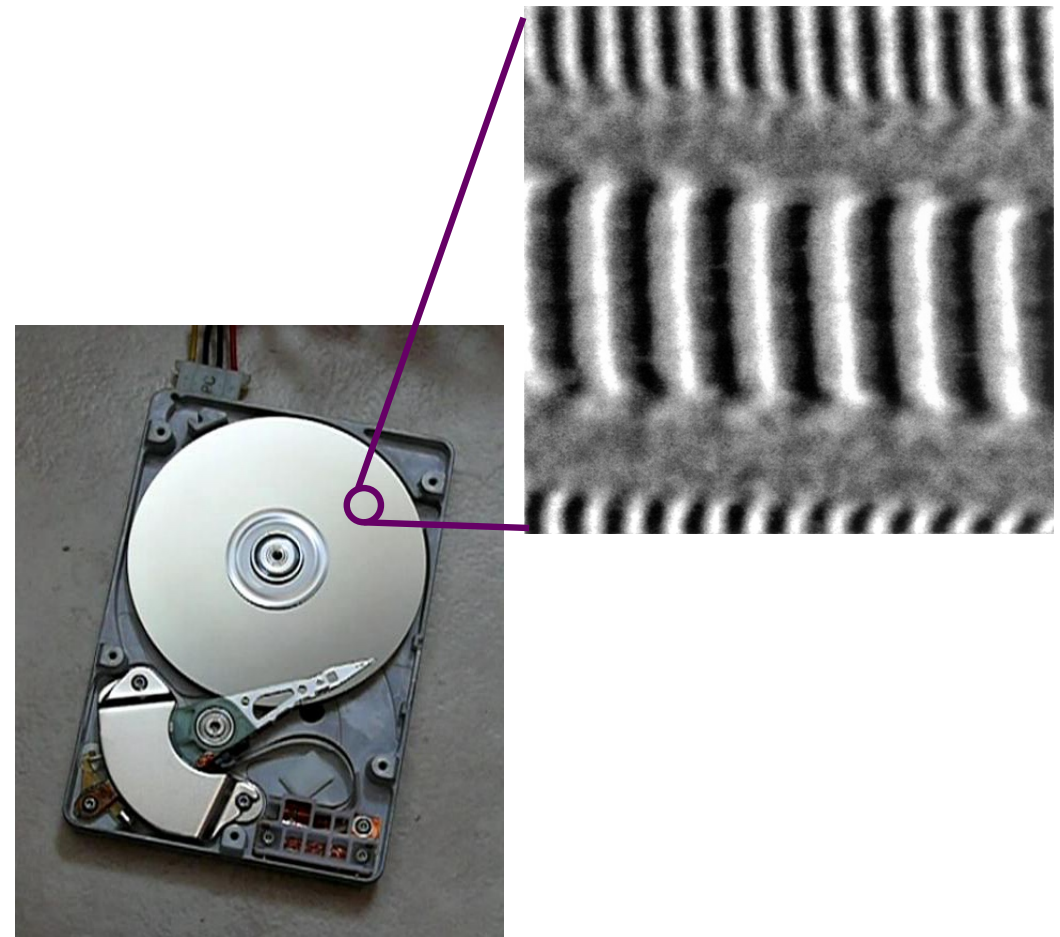
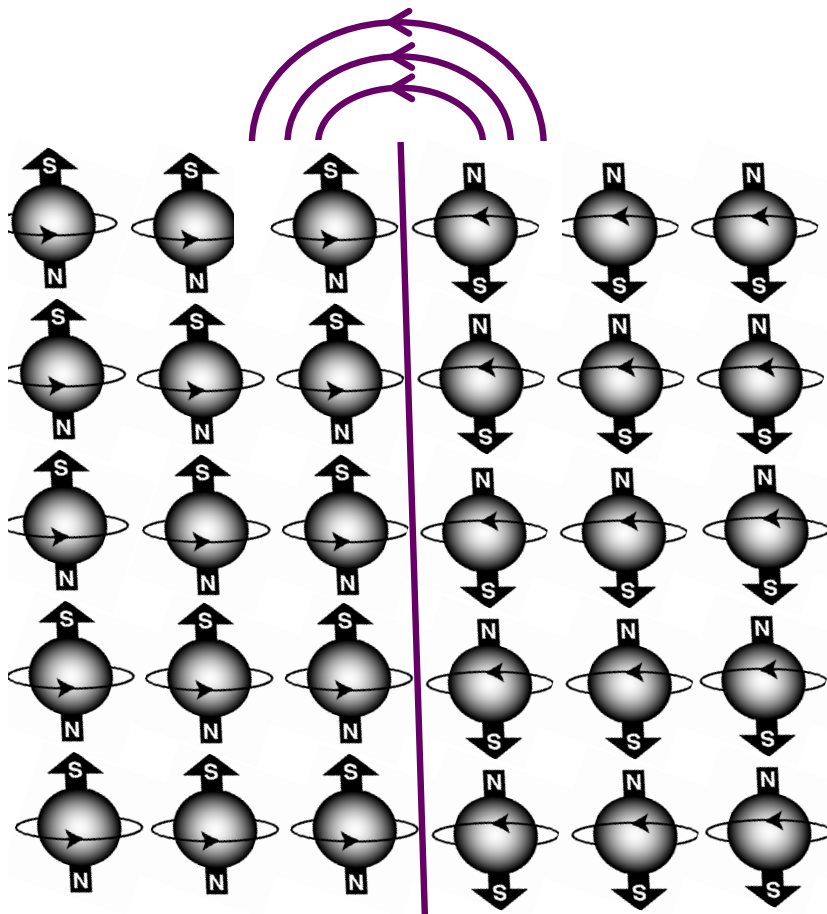
Multiferroics: Multiple ferroic orders...



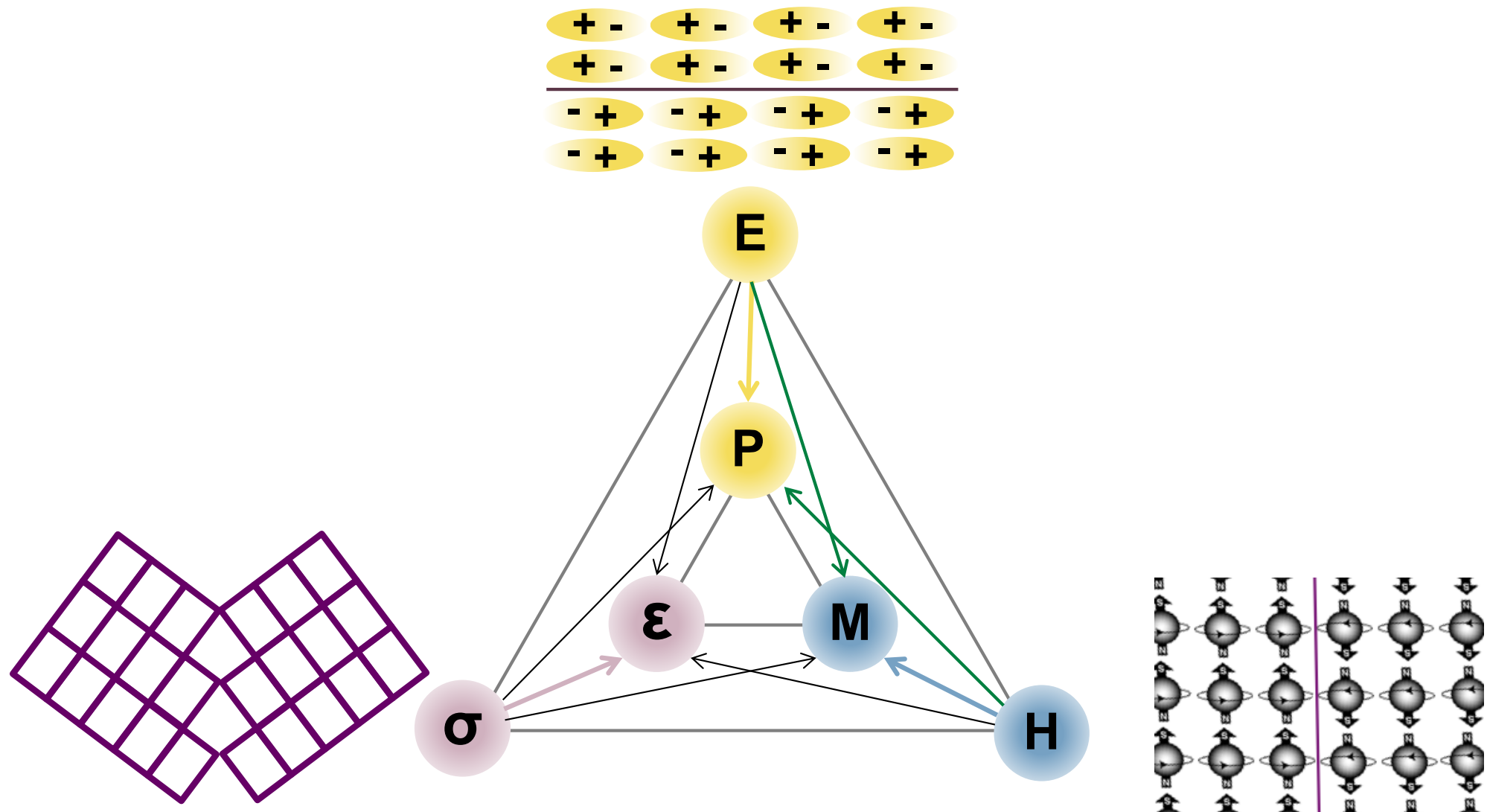
...and multiple defects from spontaneous symmetry-lowering transitions



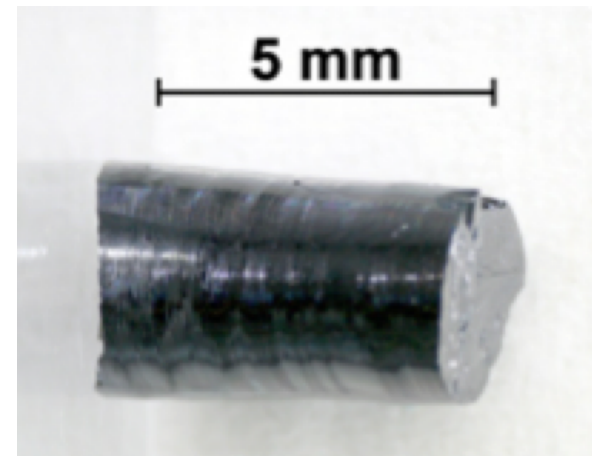
Ferromagnetic domain walls



...and multiple defects from spontaneous symmetry-lowering transitions



Our material: Multiferroic YMnO_3

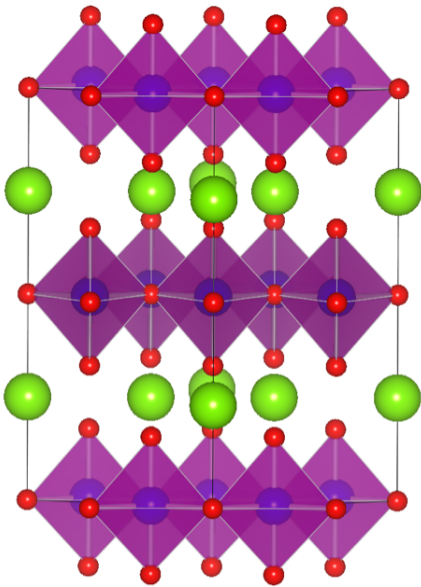


Frank Lichtenberg

Our material: Multiferroic YMnO_3

High temperature

$P6_3/mmc$

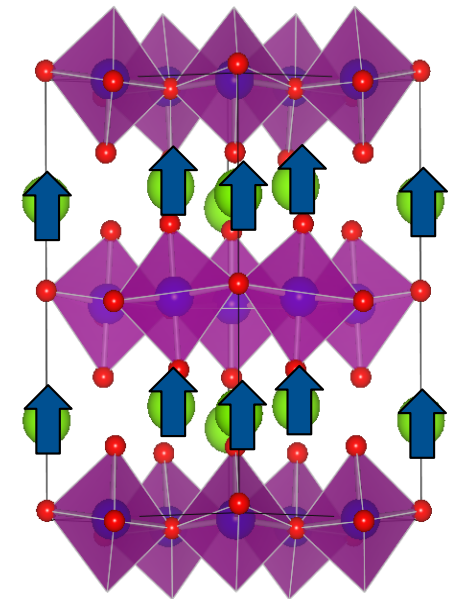


paraelectric

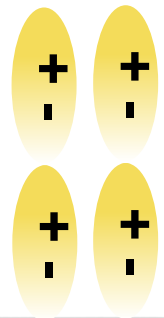
Symmetry-lowering phase
transition at $\sim 1000\text{K}$

Low temperature

$P6_3cm$



ferroelectric



Use symmetry analysis and electronic structure calculations to determine the form of the potential

Landau free energy

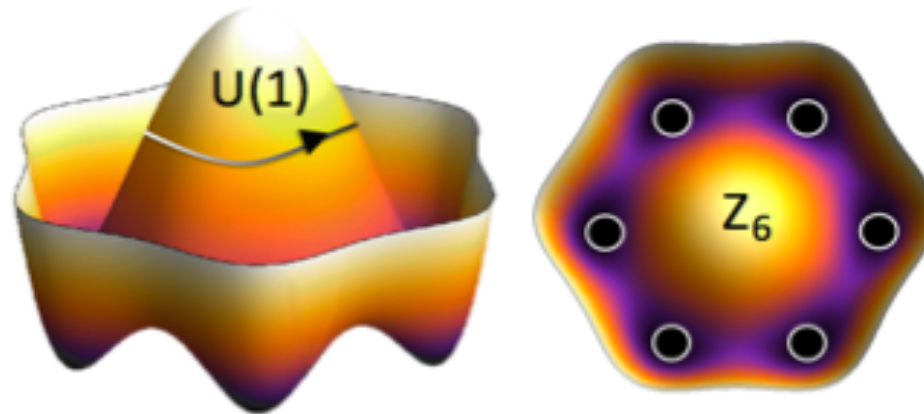
$$f_u = \frac{a}{2}Q^2 + \frac{b}{4}Q^4 + \frac{Q^6}{6}(c + c' \cos 6\Phi) - gQ^3 P_z \cos 3\Phi$$

Q is amplitude of tilting

Φ is angle of tilting

P_z is polarization

The phase transition is described by a Mexican-hat-like potential!



S. Artyukhin, K.T. Delaney, NAS and M. Mostovoy, *Landau theory of topological defects in multiferroic hexagonal manganites*, Nature Materials, 13, 42 (2014)

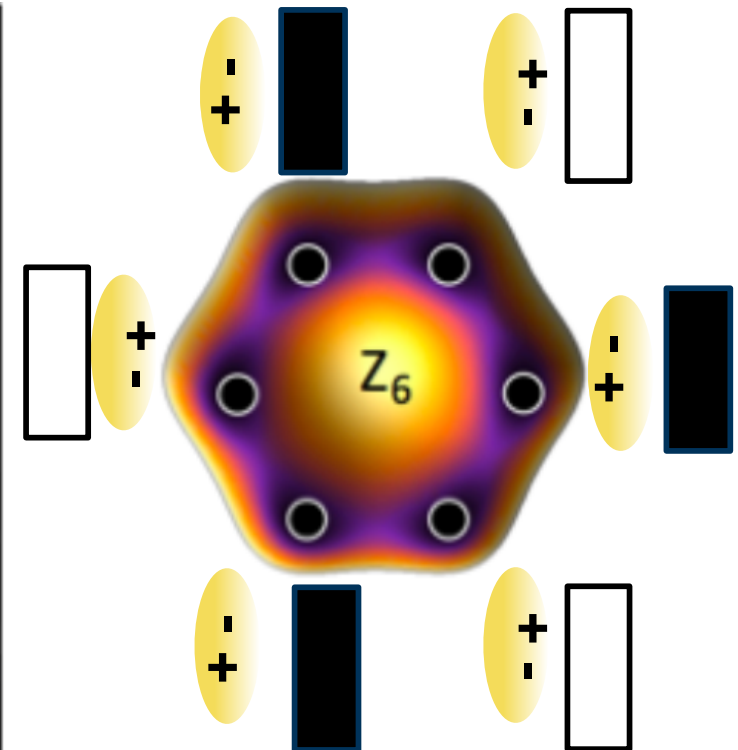
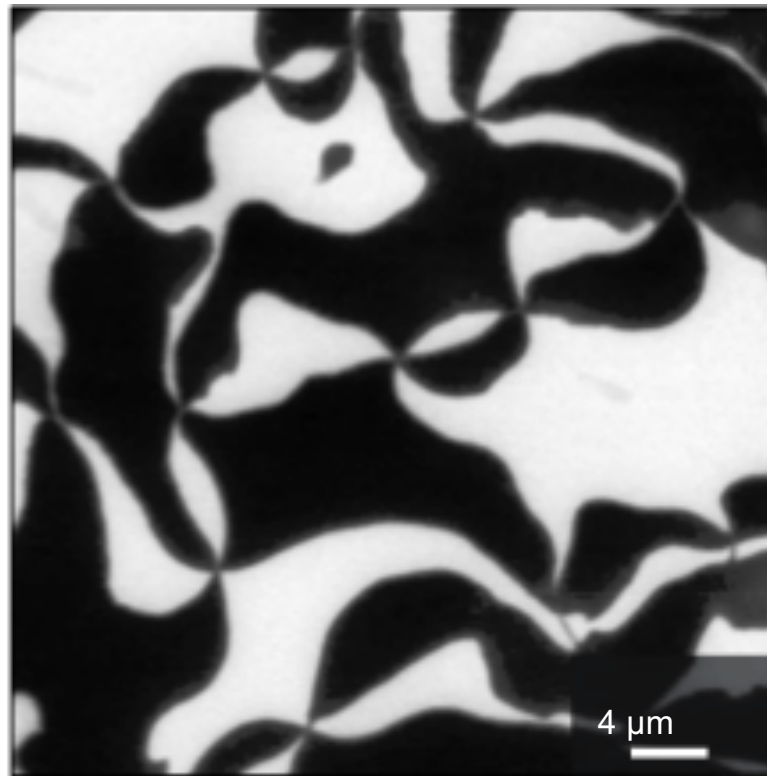
And the six minima give us six low-symmetry polar domains

Piezoforce Microscopy Image of ferroelectricity in YMnO_3

Martin
Lilienblum



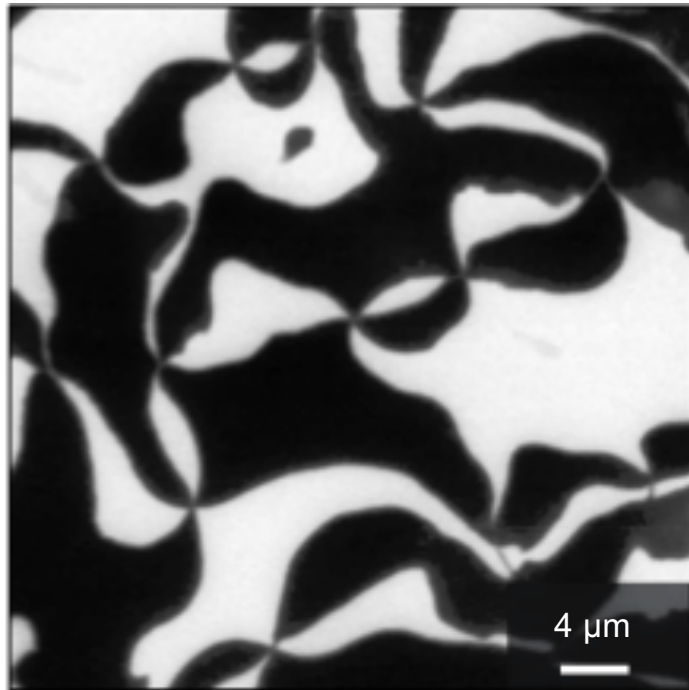
Manfred Fiebig



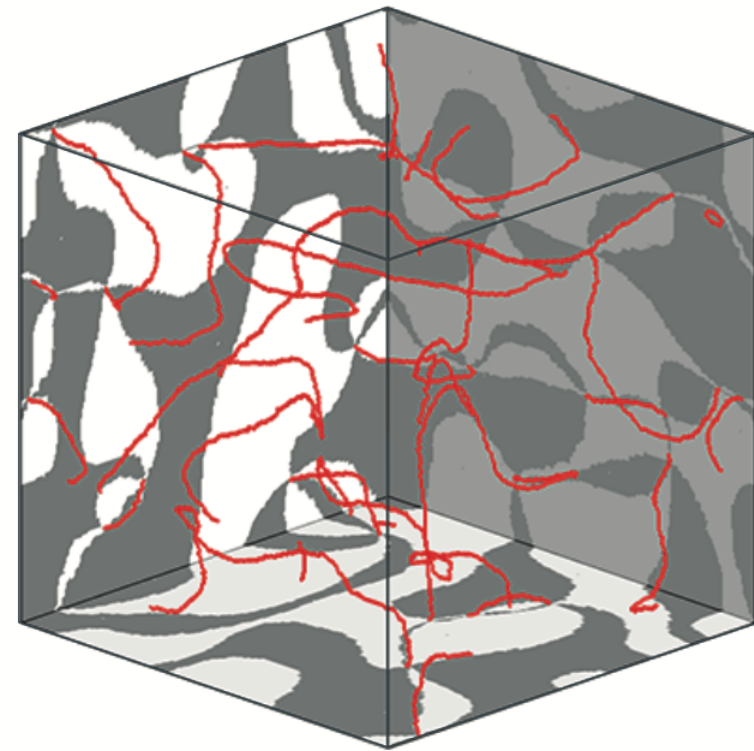
S. Artyukhin, K.T. Delaney, NAS and M. Mostovoy, *Landau theory of topological defects in multiferroic hexagonal manganites*, Nature Materials, 13, 42 (2014)

The meeting points of the ferroelectric domains are in fact one-dimensional “strings”

Piezoforce Microscopy Image of
the Defects in YMnO₃



3-D Simulation

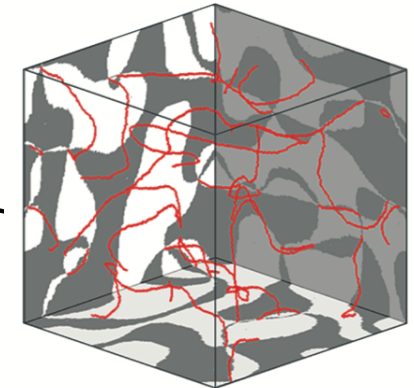
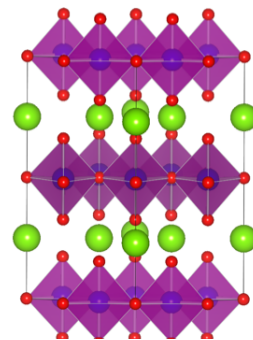
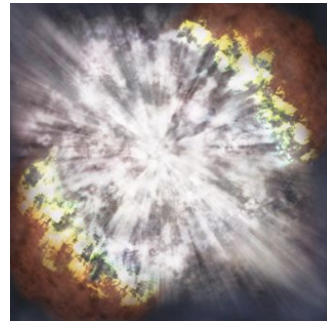
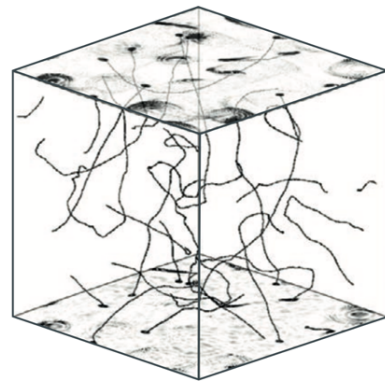


The structural phase transition in multiferroic YMnO_3 provides an analogue to the Grand Unification Transition

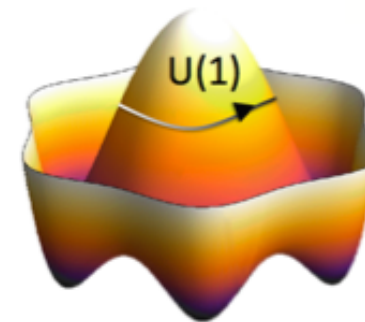
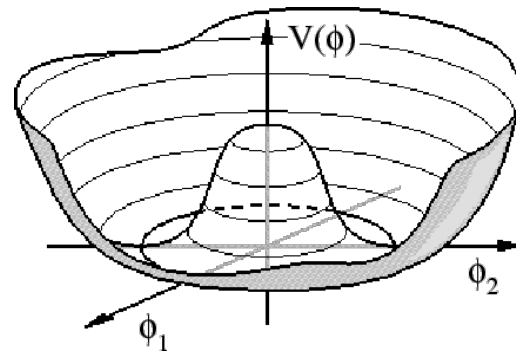
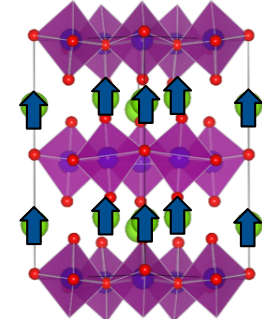
Early Universe

 YMnO_3

High symmetry vacuum

 $P6_3/mmc$ 

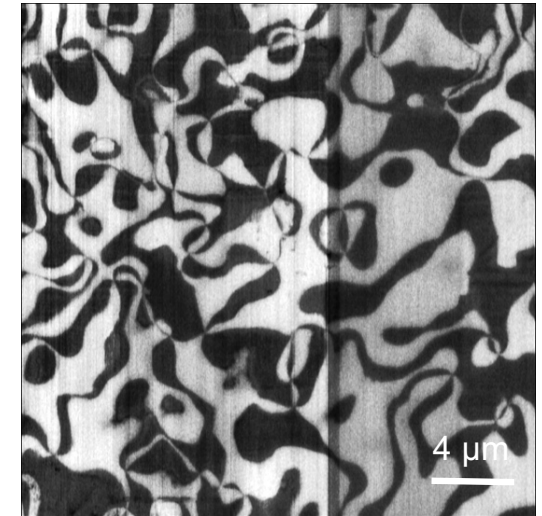
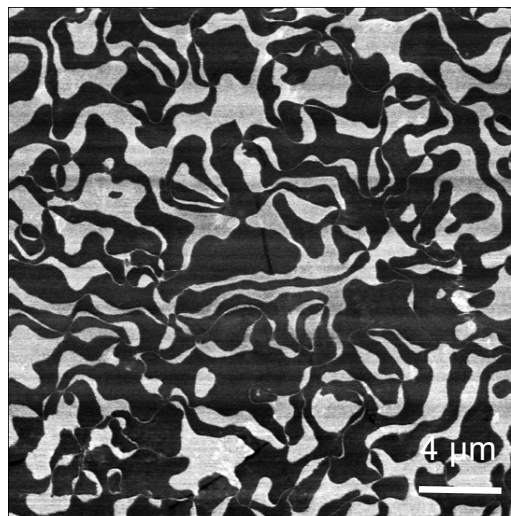
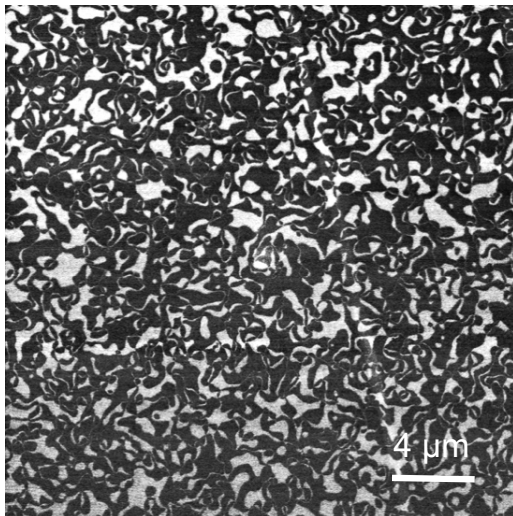
Low symmetry vacuum

 $P6_3cm$ 

What experiment would we like to do on the early universe?

We'd like to expand it at different rates, crossing the GUT, and see how many cosmic strings form in each case ("Kibble-Zurek scaling")

Instead we will cool YMnO_3 at different rates through the structural phase transition and count how many domain intersections form



Test of quantitative prediction of string formation (Kibble-Zurek scaling)

domain size,

defined as T_c ($\sim 1000\text{K}$) / cooling rate

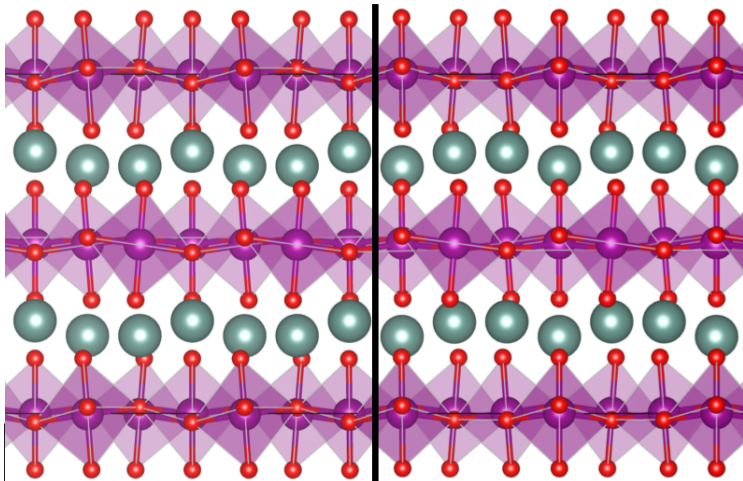
$$d = \xi_0 \left(\frac{\tau_q}{\tau_0} \right)^{\frac{\nu}{1+\mu}}$$

Critical exponents: ratio = 0.58
from MC simulations for 3D XY model

M. Campostrini et al., Phys. Rev. B **74**, 144506 (2006)

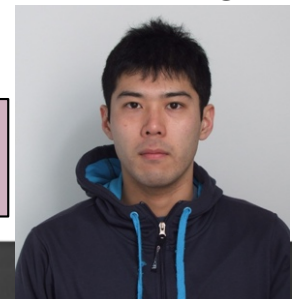
zero-temperature correlation length
~ domain wall width in ferroelectrics

zero-temperature relaxation time
= ξ_0 / speed of sound
speed of sound = 640 m/s (DFT)



Yu Kumagai and NAS, *Structural domain walls in polar hexagonal manganites*, Nat. Comm. 4, 1540 (2013)

Yu Kumagai

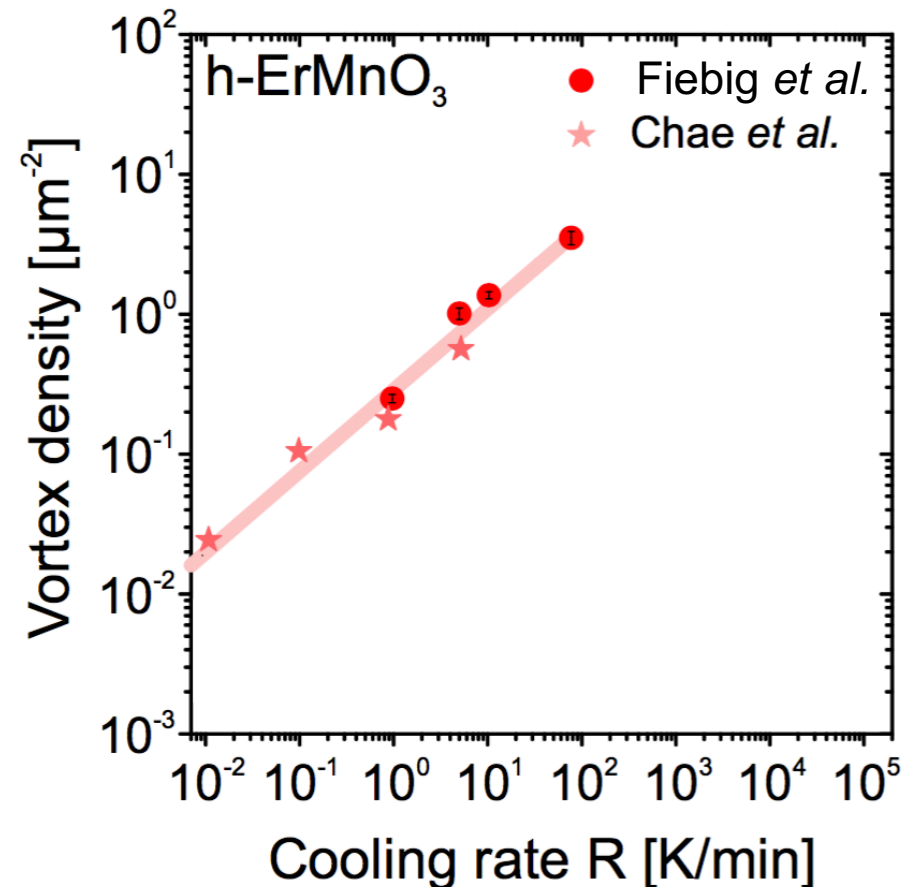


Comparison of predicted Kibble-Zurek scaling with experiment

Red line: our calculations
with $\xi_0 = 0.06 \text{ A}$

Red points: measured

REMARKABLE AGREEMENT!



S. C.. Chae *et al.*, *Direct observation of the proliferation of ferroelectric loop domains and vortex-antivortex pairs*, PRL **108**, 167603 (2012)

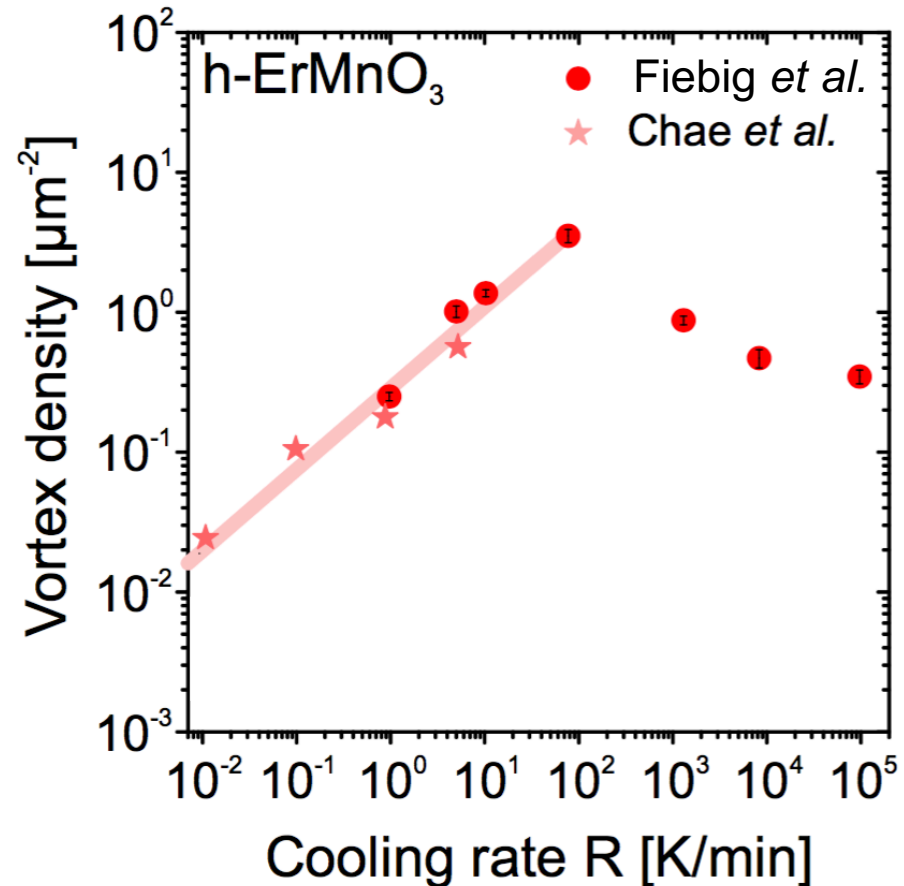
S. Griffin, M. Lilienblum, K. Delaney, Y. Kumagai, M. Fiebig and N. A. Spaldin, *Scaling behaviour and beyond equilibrium in the hexagonal manganites*, PRX **2**, 041022 (2012)

Comparison of predicted Kibble-Zurek scaling with experiment

Red line: our calculations
with $\xi_0 = 0.06 \text{ A}$

Red points: measured

REMARKABLE AGREEMENT
AT SLOW COOLING RATE!



Discovery of a “beyond-Kibble-Zurek regime” at fast cooling

Open questions:

What is the origin of the turnaround?

What is the physics of the beyond-KZ regime?

Is it relevant for early-universe behavior?



Quintin Meier

Summary

YMnO₃ seems to provide the first example of Kibble-Zurek scaling in a condensed matter system

Cosmic strings formed the way cosmologists thought ;)

The
Economist

Table-top astrophysics

How to build a multiverse

Small models of cosmic phenomena are shedding light on the real thing

Mar 16th 2013 | From the print edition

f Like

953

t

Whether all this ingenuity unravels any cosmic truth is uncertain. Cliff Burgess, a theorist at Perimeter Institute for Theoretical Physics in Ontario, has his doubts. But he thinks that such experiments are nevertheless worth pursuing. "Like tap-dancing snakes," he says, "the point is not that they do it well, it is that they do it at all."

Always go to the SERC (CSCS) annual meeting!